

(2/7/07)

①

Classical Field Theory

Lagrangian Mechanics

Action

$$S = \int dt L(q_i(t), \dot{q}_i(t))$$

$$\frac{\delta S}{\delta q_i(t)} = 0 \quad \Rightarrow \quad \underline{\text{E.O.M.}}$$

Explicitly,

$$\begin{aligned} \delta S &= \int dt \left(\frac{\partial L}{\partial q_i} \delta q_i + \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i \right) \\ &= \int dt \left(\frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} \right) \delta q_i + \text{boundary terms} = 0 \end{aligned}$$

E.O.M.:

$$\frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} = 0$$

Euler-Lagrange
Eqs.

EXAMPLE: S.H.O. 1-d

$$L = \frac{1}{2} \dot{x}^2 - \frac{\omega^2}{2} x^2$$

$$\frac{\partial L}{\partial x} - \frac{d}{dt} \frac{\partial L}{\partial \dot{x}} = -\omega^2 x - \ddot{x} = 0$$

$$\Rightarrow \underline{\underline{\ddot{x} = -\omega^2 x}}$$

Lagrangian Field Theory

Scalar field $\phi = \phi(x^\mu) =$ continuous family of D.O.F.

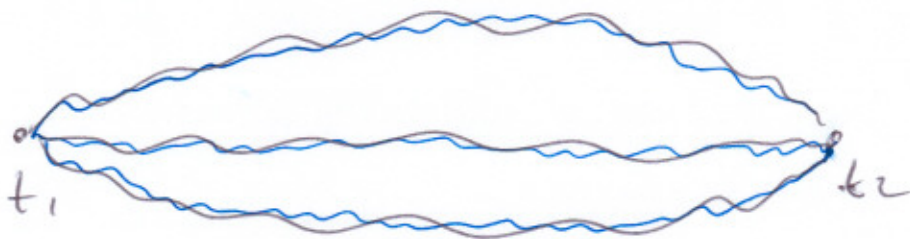
{ Can think of as limit of countable D.O.F. (as in a space-time lattice) }

Fundamental quantity is S : the action

$$S = \int L dt = \int \mathcal{L}(\phi, \partial_\mu \phi) d^4x$$

\mathcal{L} is Lagrange density

"Principle of least action."



When system evolves from one configuration to another between t_1 and t_2 it does so along path for which S is an extremum.
(usually minimum)

$$0 = \delta S$$

$$= \int d^4x \left(\frac{\partial \mathcal{L}}{\partial \phi} \delta \phi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta (\partial_\mu \phi) \right)$$

$$= \int d^4x \left(\frac{\partial \mathcal{L}}{\partial \phi} \delta \phi - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right) \delta \phi + \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \delta \phi \right) \right)$$

Surface integral over S.T. boundary

A) $\delta\phi$ is 0 at temporal boundaries as have initial + final field configurations.

B) Restrict ourselves to $\delta\phi$ that vanish on spatial boundary



Surface term vanishes

As $\delta\phi$ is arbitrary,

$$\partial_\mu \left(\frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \right) - \frac{\partial \mathcal{L}}{\partial \phi} = 0$$

Euler-Lagrange
E.O.M.

(1 such equation for each field.)

EXAMPLE: $U(1)$ field

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2$$

$$\frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} = \partial^\mu \phi$$

$$\frac{\partial \mathcal{L}}{\partial \phi} = -m^2 \phi$$

E.O.M. KLEIN-GORDON EQUATION

$$\partial_\mu \partial^\mu \phi + m^2 \phi = 0$$

or

$$(\square + m^2) \phi = 0$$

(Relativistic version of S.H.O. equation)

Hamiltonian Mechanics

$$p_i = \frac{\partial L}{\partial \dot{q}_i} \quad \{q_i, p_j\} = \delta_{ij}$$

$$H = \sum_i p_i \dot{q}_i - L$$

Hamiltonian Field Theory

Lagrangian formulation well suited to relativity, as is manifestly Lorentz invariant.

Hamiltonian formulation is more intuitive as related to non-relativistic C.M.

But

$$H \sim E \rightarrow E' \neq E$$

Boosts

LOSE Manifest Lorentz invariance!

Nevertheless we will follow the Hamiltonian method.

By analogy w/ mechanics, assume

$$\pi(x^i) = \frac{\partial \mathcal{L}}{\partial \dot{\phi}(x^i)}$$

$$H = \int d^3x \mathcal{H} = \int d^3x [\pi(x^i) \dot{\phi}(x^i) - \mathcal{L}]$$

\mathcal{H} is Hamiltonian density

EXAMPLE: K-G field

$$\begin{aligned} \mathcal{L} &= \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 \\ &= \frac{1}{2} (\partial_\mu \phi)^2 - \frac{1}{2} m^2 \phi^2 \end{aligned}$$

$$\Rightarrow \mathcal{L} = \frac{1}{2} \dot{\varphi}^2 - \frac{1}{2} (\nabla_i \varphi)^2 - \frac{1}{2} m^2 \varphi^2$$

$$\pi(x^i) = \frac{\partial \mathcal{L}}{\partial \dot{\varphi}} = \dot{\varphi}(x^i)$$

$$H = \int d^3x \left[\dot{\varphi}^2 - \frac{1}{2} \dot{\varphi}^2 + \frac{1}{2} (\nabla_i \varphi)^2 + \frac{1}{2} m^2 \varphi^2 \right]$$

↓

$$H = \int d^3x \left[\frac{1}{2} \dot{\varphi}^2 + \frac{1}{2} (\nabla_i \varphi)^2 + \frac{1}{2} m^2 \varphi^2 \right]$$

$$H = \int d^3x \left[\frac{1}{2} \pi^2 + \frac{1}{2} (\nabla_i \varphi)^2 + \frac{1}{2} m^2 \varphi^2 \right]$$

Energy cost:

moving in
time

moving in
space

existing at
all

Noether's Theorem

Relation between symmetries + conservation laws

Consider infinitesimal transformations on φ .

(*)
$$\varphi(x) \rightarrow \varphi'(x) = \varphi(x) + \alpha \Delta \varphi(x)$$

\uparrow
 infinitesimal
parameter

\uparrow
 Deformation
of
field

This transformation is Symmetry if the $E-L$ E.O.M. are left invariant.

\updownarrow
 \int invariant

Hence, w/ respect to (*),

$$\mathcal{L}(x) \rightarrow \mathcal{L}(x) + \alpha \partial_\mu J^\mu(x)$$

\uparrow
 surface term,
 (vanishes at boundary of S.T.)

($J^\mu(x)$ is fn. of fields)

Compare this with direct variation of \mathcal{L}

$$\mathcal{L}(x) \rightarrow \mathcal{L}(x) + \alpha \Delta \mathcal{L}$$

$$\begin{aligned}
\alpha \Delta \mathcal{L} &= \alpha \left(\frac{\partial \mathcal{L}}{\partial \varphi} \Delta \varphi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta (\partial_\mu \varphi) \right) \\
&= \frac{\partial \mathcal{L}}{\partial \varphi} (\alpha \Delta \varphi) + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \partial_\mu (\alpha \Delta \varphi) \\
&= \alpha \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi \right) + \alpha \underbrace{\left[\frac{\partial \mathcal{L}}{\partial \varphi} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \right) \right]}_{\text{E-L E.O.M.} \Rightarrow 0} \Delta \varphi
\end{aligned}$$

Comparing w/ previous \Rightarrow

$$\partial_\mu J^\mu(x) = \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi \right)$$

\Downarrow solution

$$J^\mu(x) \equiv \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi - j^\mu(x)$$

where $\partial_\mu j^\mu(x) = 0$

"current" $j^\mu(x)$ is conserved!

(Invert: $j^\mu = \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \Delta \varphi - J^\mu$)

∴ For each continuous symmetry of \mathcal{L}
 \Downarrow
 conservation Law

(Noether's Theorem)

(Clear generalization to wave fields)

Note:

Conservation law \Rightarrow conserved charge

$$Q \equiv \int j^0(x) d^3x$$

$$\frac{\partial Q}{\partial t} = \int \partial_0 j^0(x) d^3x$$

$$= \int \partial_\mu j^\mu(x) d^3x + \text{surface terms}$$

$$\Rightarrow \boxed{\frac{\partial Q}{\partial t} = 0}$$

(Not in as solid ground!)

EXAMPLES:

(A)

$$\mathcal{L} = \frac{1}{2} (\partial_\mu \phi)^2$$

Invariance: $\phi \rightarrow \phi + \alpha$ α constant

Recall:

$$j^\mu(x) = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} \Delta \phi - \cancel{j^\mu}$$

(not relevant here)

$$j^\mu = \partial^\mu \phi \quad \partial_\mu j^\mu = \square \phi = 0$$

(K-G equation)

(B)

$$\mathcal{L} = |\partial_\mu \phi|^2 - m^2 |\phi|^2$$
$$= \partial_\mu \phi \partial^\mu \phi^* - m^2 \phi \phi^*$$

ϕ is complex valued field

K-G equation?

treat ϕ, ϕ^* as independent fields

$$\varphi: \partial_\mu (\partial^\mu \varphi^\dagger) + m^2 \varphi^\dagger = 0 \Rightarrow (\mathcal{D} + m^2) \varphi^\dagger = 0$$

$$\varphi^\dagger: (\mathcal{D} + m^2) \varphi = 0$$

Invariance of \mathcal{L} ?

$\varphi \rightarrow e^{i\alpha} \varphi$

"phase transformation"

infinitesimal: $\varphi \rightarrow \varphi + i\alpha \varphi + \mathcal{O}(\alpha^2)$

$$\begin{aligned} \Delta \varphi &= i\alpha \varphi \\ \Delta \varphi^\dagger &= -i\alpha \varphi^\dagger \end{aligned}$$

Conserved current?

$$j^\mu(x) = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \varphi)} \Delta \varphi + \frac{\partial \mathcal{L}}{\partial(\partial_\mu \varphi^\dagger)} \Delta \varphi^\dagger$$

$$= (\partial^\mu \varphi^\dagger)(i\varphi) + (\partial^\mu \varphi)(-i\varphi^\dagger)$$

\Rightarrow $j^\mu(x) = i [(\partial^\mu \varphi^\dagger) \varphi - \varphi^\dagger \partial^\mu \varphi]$

Conserved?

$$\partial_\mu j^\mu = i \left[(\mathcal{D}\varphi^\dagger)\varphi + \partial^\mu \varphi^\dagger \partial_\mu \varphi - \partial_\mu \varphi^\dagger \partial^\mu \varphi - \varphi^\dagger (\mathcal{D}\varphi) \right]$$

$$\stackrel{\text{(use K-6)}}{=} i \left[-m^2 |\varphi|^2 + m^2 |\varphi|^2 \right] = 0 \quad \blacktriangle$$

Noether's Thm also applies to

Space-time transformations

Consider S-T translations:

$$\underline{x^\mu \rightarrow x^\mu - a^\mu} \quad (a^\mu \text{ is constant})$$

$$\varphi: \quad \varphi(x) \rightarrow \varphi(x+a) = \varphi(x) + a^\mu \partial_\mu \varphi(x) + \mathcal{O}(a^2)$$

$$\psi: \quad \psi \rightarrow \psi + a^\mu \partial_\mu \psi = \psi + a^\nu \partial_\mu (\delta^\mu_\nu \psi)$$

4 separately conserved currents:

$$T^\mu_\nu \equiv \frac{\partial \mathcal{L}}{\partial (\partial_\mu \varphi)} \partial_\nu \varphi - \mathcal{L} \delta^\mu_\nu$$

T^{μ}_{ν} : Energy - Momentum Tensor

Let's look more carefully at δS w.r. respect to S-T translations:

$$\mathcal{L} = \mathcal{L}(\varphi, \partial_{\mu}\varphi, x^{\mu})$$

$$\underline{x^{\mu} \rightarrow x^{\mu'} = x^{\mu} + \delta x^{\mu}}$$

$$\delta S = \int d^4x' \mathcal{L}(\varphi', \partial_{\nu}\varphi', x^{\mu'}) - \int d^4x \mathcal{L}(\varphi, \partial_{\nu}\varphi, x^{\mu})$$

$$\underline{d^4x' = J(\underline{x}') d^4x} \quad (J \text{ is Jacobian})$$

$$\frac{\partial x'^{\mu}}{\partial x^{\lambda}} = \delta^{\mu}_{\lambda} + \partial_{\lambda} \delta x^{\mu}$$

$$\underline{J(\underline{x}') = \det\left(\frac{\partial x'^{\mu}}{\partial x^{\lambda}}\right) = 1 + \partial_{\mu}(\delta x^{\mu})}$$

$$\Rightarrow \delta S = \int d^4x (\delta \mathcal{L} + \mathcal{L} \partial_{\mu} \delta x^{\mu})$$

$$\left(\delta \mathcal{L} = \frac{\partial \mathcal{L}}{\partial \varphi} \delta \varphi + \frac{\partial \mathcal{L}}{\partial(\partial_{\nu}\varphi)} \delta(\partial_{\nu}\varphi) + \frac{\partial \mathcal{L}}{\partial x^{\mu}} \delta x^{\mu} \right)$$

$$S_0 \quad \delta S = \int d^4x \left(\frac{\partial \mathcal{L}}{\partial \psi} \delta \psi + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi)} \delta (\partial_\mu \psi) + \partial_\mu (\mathcal{L} \delta x^\mu) \right)$$

Source of extra piece!

Conservation Laws

$$\begin{aligned} T^{00} &= \frac{\partial \mathcal{L}}{\partial \dot{\psi}} \dot{\psi} - \mathcal{L} \\ &= \pi(x) \dot{\psi}(x) - \mathcal{L} = \mathcal{H} \end{aligned}$$

$$S_0 \quad H = \int d^3x T^{00}$$

H is generator of time translations

$$T^0_i = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} \partial_i \psi = \pi \partial_i \psi$$

$$S_0 \quad P^i = \int d^3x T^{0i} = - \int d^3x \pi \partial_i \psi$$

Momenta (carried by field) generator of space translations